VELOCITY CUMULANTS AND CORRELATIONS IN A GRANULAR GAS OF ROUGH SPHERES

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Una manera de hacer Europa



- It is a conglomeration of discrete solid, macroscopic particles characterized by a loss of energy whenever the grains collide.
- The constituents must be large enough such that they are not subject to thermal motion fluctuations. Thus, the lower size limit for grains is about 1 µm.

• Examples of granular materials would include nuts, coal, sand, rice, coffee, corn flakes, fertilizer, ball bearings, ...



• ... and even Saturn's rings



- Granular materials are commercially important in applications as diverse as pharmaceutical industry, agriculture, and energy production.
- They are ubiquitous in nature and are the second-most manipulated material in industry (the first one is water).





WHAT IS A GRANULAR *FLUID*?

• When the granular matter is driven and energy is fed into the system (e.g., by shaking) such that the grains are not in constant contact with each other, the granular material is said to *fluidize*.



Experiment (A. Kudrolli's group)

Simulations (D.C. Rapaport)



SIMPLE MODEL OF A GRANULAR GAS: A *COLLECTION* OF *INELASTIC ROUGH* HARD SPHERES

This model unveils an inherent breakdown of energy equipartition in granular fluids, even in homogeneous and isotropic states





OUTLINE OF THE TALK

- Collision rules for inelastic rough hard spheres. Statistical quantities.
- Homogeneous cooling state. Kinetic theory (Boltzmann-Enskog) description.
- Sonine approximation. Results.
- Conclusions and outlook.

MATERIAL PARAMETERS:

- Mass *m*
- Diameter *o*
- Moment of inertia *I*
- Coefficient of normal restitution α
- Coefficients of tangential restitution β
- $\alpha = 1$ for perfectly elastic particles
- β =-1 for perfectly smooth particles
- $\beta = +1$ for perfectly rough particles

Collision rules

 ω_i

 $\frac{\sigma}{2}\widehat{\boldsymbol{\sigma}}$

 \mathbf{v}_{ij}

Cons. linear momentum: $\mathbf{v}'_i + \mathbf{v}'_j = \mathbf{v}_i + \mathbf{v}_j$

Cons. angular momentum:

 $egin{aligned} &Im{\omega}_{i,j}' \mp mrac{\sigma_i}{2}\widehat{m{\sigma}} imes \mathbf{v}_{i,j}' \ &= Im{\omega}_{i,j} \mp mrac{\sigma_i}{2}\widehat{m{\sigma}} imes \mathbf{v}_i \end{aligned}$

Relative velocity of the points of the spheres at contact:

$$oldsymbol{\mathcal{V}}_{ij} = \mathbf{v}_{ij} - rac{\sigma}{2} \widehat{oldsymbol{\sigma}} imes (oldsymbol{\omega}_i + oldsymbol{\omega}_j)$$

$$\left| \widehat{\boldsymbol{\sigma}} \cdot \boldsymbol{\mathcal{V}}_{ij}' = -lpha \widehat{\boldsymbol{\sigma}} \cdot \boldsymbol{\mathcal{V}}_{ij}, \quad \widehat{\boldsymbol{\sigma}} imes \boldsymbol{\mathcal{V}}_{ij}' = -eta \widehat{\boldsymbol{\sigma}} imes \boldsymbol{\mathcal{V}}_{ij}
ight|$$

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 $\frac{\sigma}{2}\widehat{\boldsymbol{\sigma}}$

ENERGY COLLISIONAL LOSS

$$E_{ij} = \frac{1}{2}mv_i^2 + \frac{1}{2}mv_j^2 + \frac{1}{2}I\omega_i^2 + \frac{1}{2}I\omega_j^2$$
$$E'_{ij} = -(1 - \alpha^2) \times \cdots$$

$$E'_{ij} - E_{ij} = -(1 - \alpha^2) \times \cdots$$

 $-(1 - \beta^2) \times \cdots$

Energy is conserved *only* if the spheres are

- elastic (α =1) and
- either
 - perfectly smooth (β =-1) or
 - perfectly rough (β =+1)



GRANULAR TEMPERATURES, KURTOSES, AND CORRELATIONS



translational temperature: $\langle v^2 \rangle = \frac{3T^{\rm tr}}{-}$ rotational temperature: $\langle \omega^2 \rangle = \frac{3T^{\rm rot}}{r}$ translational kurtosis: $\langle v^4 \rangle = \frac{5}{3} \langle v^2 \rangle^2 (1 + a_{20})$ rotational kurtosis: $\langle \omega^4 \rangle = \frac{5}{3} \langle \omega^2 \rangle^2 (1 + a_{02})$ scalar correlations: $\langle v^2 \omega^2 \rangle = \langle v^2 \rangle \langle \omega^2 \rangle (1 + a_{11})$ angular correlations: $\langle (\widehat{\mathbf{v}} \cdot \widehat{\boldsymbol{\omega}})^2 \rangle = \frac{1}{3} \left(1 + \frac{3b}{5} \right)$

OUR AIM:

To measure

- Temperature ratio $T^{\rm rot}/T^{\rm tr}$
- Kurtosis a_{20}
- Kurtosis a_{02}
- Correlation a_{11}
- Correlation b

in the Homogeneous Cooling State (HCS).

$$T^{\mathrm{tr}}(t) \sim t^{-2}, \quad T^{\mathrm{rot}}(t)/T^{\mathrm{tr}}(t) \to \mathrm{const}$$





David Enskog (1884-1947)



Linking Bully

(1844-1906)

(Cartoon by Bernhard) Reischl, University of Vienna)

Boltzmann-Enskog equation: $\partial_t f(\mathbf{r}, \mathbf{v}, \boldsymbol{\omega}, t) + \mathbf{v} \cdot \nabla f(\mathbf{r}, \mathbf{v}, \boldsymbol{\omega}, t) = J[\mathbf{r}, \mathbf{v}, \boldsymbol{\omega}, t|f]$

Inelastic+Rough collisions

SCALED QUANTITIES

Scaled velocities: $\mathbf{c}(t) \equiv \frac{\mathbf{v}}{\sqrt{2T^{\mathrm{tr}}(t)/m}}, \quad \mathbf{w}(t) \equiv \frac{\omega}{\sqrt{2T^{\mathrm{rot}}(t)/I}}$ 1 $\left[AT^{\mathrm{tr}}(t)T^{\mathrm{rot}}(t) \right]^{3/2}$

Scaled distribution function: $\phi(\mathbf{c}, \mathbf{w}) \equiv \frac{1}{n} \left[\frac{4T^{\mathrm{tr}}(t)T^{\mathrm{rot}}(t)}{mI} \right]^{3/2} f(\mathbf{v}, \boldsymbol{\omega}, t)$

HCS:
$$\frac{\mu_{20}}{3} \frac{\partial}{\partial \mathbf{c}} \cdot (\mathbf{c}\phi) + \frac{\mu_{02}}{3} \frac{\partial}{\partial \mathbf{w}} \cdot (\mathbf{w}\phi) = J^*[\mathbf{c}, \mathbf{w}|\phi]$$

Collisional moments:

$$\mu_{pq} = -\int d\mathbf{c} \int d\mathbf{w} \, c^p w^q J^*[\mathbf{c}, \mathbf{w} | \phi]$$
$$\mu_b = -\int d\mathbf{c} \int d\mathbf{w} \, (\mathbf{c} \cdot \mathbf{w})^2 J^*[\mathbf{c}, \mathbf{w} | \phi]$$

MOMENT EQUATIONS

 $\mu_{20} = \mu_{02}$

$$5\mu_{20} = \frac{\mu_{40}}{1+a_{20}}$$

$$\frac{3}{2}(\mu_{20} + \mu_{02}) = \frac{\mu_{22}}{1 + a_{11}}$$

$$5\mu_{02} = \frac{\mu_{04}}{1+a_{02}}$$

$$rac{5}{6}(\mu_{20}+\mu_{02})b=\mu_b-rac{1}{3}\mu_{22}$$

LINEAR SONINE APPROXIMATION

$$\begin{split} \phi(\mathbf{c}, \mathbf{w}) &\simeq \pi^{-3} e^{-c^2 - w^2} \left\{ 1 + a_{20} S_{\frac{1}{2}}^{(2)}(c^2) + a_{02} S_{\frac{1}{2}}^{(2)}(w^2) \right. \\ &\left. + a_{11} S_{\frac{1}{2}}^{(1)}(c^2) S_{\frac{1}{2}}^{(1)}(w^2) + b \left[(\mathbf{c} \cdot \mathbf{w})^2 - \frac{1}{3} c^2 w^2 \right] \right\} \end{split}$$

Sonine (Laguerre) polynomials: $S_{\frac{1}{2}}^{(1)}(x) = \frac{3}{2} - x$, $S_{\frac{1}{2}}^{(2)}(x) = \frac{1}{8} \left(15 - 20x + 4x^2\right)$

AND AFTER TEDIOUS CALCULATIONS ...

$$\begin{split} \mu_{20} &= 4\sqrt{2\pi} \left[\left(\tilde{\alpha}(1-\tilde{\alpha}) + \tilde{\beta}(1-\tilde{\beta}) \left(1 + \frac{3a_{00}}{16} \right) - \theta \frac{\beta^2}{\kappa} \left(1 - \frac{a_{20}}{16} + \frac{3a_{11}-b}{12} \right) \right], \\ \mu_{02} &= 4\sqrt{2\pi} \frac{\beta}{\kappa} \left[\left(1 - \frac{\beta}{\kappa} \right) \left(1 - \frac{a_{20}}{16} + \frac{3a_{11}-b}{12} \right) - \frac{\beta}{\theta} \left(1 + \frac{3a_{01}}{16} \right) \right], \\ \mu_{40} &= 16\sqrt{2\pi} \left\{ \tilde{\alpha}^3(2-\tilde{\alpha}) + \tilde{\beta}^3(2-\tilde{\beta}) - \tilde{\alpha}\tilde{\beta}(1-\tilde{\alpha} - \tilde{\beta} + \tilde{\alpha}\tilde{\beta}) + \frac{11}{8} (\tilde{\alpha} + \tilde{\beta}) - \frac{19}{8} (\tilde{\alpha}^2 + \tilde{\beta}^2) - \left[\tilde{\alpha}\tilde{\beta} \left(\frac{23}{15} - \tilde{\alpha} - \tilde{\beta} + \tilde{\alpha}\tilde{\beta} \right) - \frac{269}{120} (\tilde{\alpha} + \tilde{\beta}) + \frac{357}{120} (\tilde{\alpha}^2 + \tilde{\beta}^2) - \tilde{\alpha}^3(2-\tilde{\alpha}) - \tilde{\beta}^3(2-\tilde{\beta}) \right] \frac{15a_{2}}{16} \\ &- \frac{11\tilde{\beta}^2 \theta}{8\kappa} \left(1 + \frac{41a_{20}}{176} + \frac{3a_{11}-b}{4} \right) + \frac{\tilde{\beta}^2 \theta}{\kappa} \left[\tilde{\alpha}(1-\tilde{\alpha}) + 2\tilde{\beta}(1-\tilde{\beta}) \right] \left(1 + \frac{3a_{20}}{36} + \frac{3a_{11}-b}{4} \right) - \frac{\tilde{\beta}^4 \theta^2}{16} \left(1 - \frac{a_{20}}{16} + \frac{3a_{11}-b}{4} \right) + \frac{\tilde{\beta}^2 \theta}{3\kappa} \left[\tilde{\alpha} - \tilde{\alpha} - \tilde{\beta} + \tilde{\alpha}\tilde{\beta} \right] \left(1 + \frac{3a_{20}}{16} + \frac{3a_{11}-b}{4} \right) + \frac{\tilde{\beta}}{3\kappa} \left(1 - \tilde{\alpha} \right) - \frac{\tilde{\beta}^3 \theta^2}{3\kappa} \left[3 - \tilde{\alpha} - \frac{\tilde{\beta}}{\kappa} + 2\frac{\tilde{\beta}^2}{\kappa} \right] \right] \left(1 + \frac{3a_{20}}{16} + \frac{3a_{11}-b}{4} \right) + \frac{\tilde{\beta}^2 \theta}{3\kappa} \left[1 - \tilde{\beta} \right] \left(1 - \tilde{\beta} \right) - \frac{8\tilde{\beta}^2 \theta^2}{3\kappa} \left(\frac{3}{4} - \tilde{\beta} - \frac{\tilde{\beta}}{\kappa} + 2\frac{\tilde{\beta}^2}{\kappa} \right) \right] \left(1 + \frac{3a_{20}}{16} + \frac{3a_{11}-b}{4} \right) + \frac{\tilde{\gamma}\tilde{\beta}}{3\kappa} \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \left(1 - \tilde{\beta} \right) - \frac{\tilde{\beta}\tilde{\beta}^2 \theta^2}{3\kappa} \left[1 - \tilde{\beta} \right] \right] \left(1 + \frac{3a_{20}}{3\kappa} + \frac{3a_{11}-b}{4} \right) + \frac{\tilde{\gamma}\tilde{\beta}}{3\kappa} \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \left(1 + \frac{\tilde{\beta}}{3\kappa} \right) \left(1 - \tilde{\beta} \right) \right] \left(1 + \frac{\tilde{\beta}\tilde{\beta}}{3\kappa} \right) \left(1 - \tilde{\beta} \right) \left(1 + \frac{\tilde{\beta}\tilde{\beta}}{3\kappa} \right) \left(1 - \tilde{\beta} \right) \right) \left(1 + \frac{\tilde{\beta}\tilde{\beta}}{3\kappa} \right) \left(1 - \tilde{\beta} \right) \right) \left(1 + \frac{\tilde{\beta}\tilde{\beta}}{3\kappa} \right) \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \right) \left(1 + \frac{\tilde{\beta}\tilde{\beta}}{3\kappa} \right) \left(1 - \frac{\tilde{\beta}}{3\kappa} \right) \right) \left(1 + \frac{\tilde{\beta}\tilde{\beta}}{3\kappa} \right) \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \right) \left(1 - \frac{\tilde{\beta}\tilde{\beta}}{3\kappa} \right) \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \right) \left(1 - \frac{\tilde{\beta}}{\kappa} \right) \left(1 - \frac{\tilde{\beta$$

INSERTING INTO THE MOMENT EQUATIONS

 $\mu_{20} = \mu_{02}$

$$5\mu_{20} = \frac{\mu_{40}}{1+a_{20}}$$

$$\frac{3}{2}(\mu_{20} + \mu_{02}) = \frac{\mu_{22}}{1 + a_{11}}$$

$$5\mu_{02} = \frac{\mu_{04}}{1+a_{02}}$$

And neglecting terms nonlinear in a_{20} , a_{11} , a_{02} and b, we get a polynomial equation for $T^{\text{rot}}/T^{\text{tr}}$ and linear equations for a_{20} , a_{11} , a_{02} and b.

$$\frac{5}{6}(\mu_{20}+\mu_{02})b=\mu_b-\frac{1}{3}\mu_{22}$$





RESULTS: KURTOSES & SCALAR CORRELATIONS





$$\langle v^4 \rangle = \frac{5}{3} \langle v^2 \rangle^2 (1 + a_{20})$$

$$\langle \omega^4 \rangle = \frac{5}{3} \langle \omega^2 \rangle^2 (1 + a_{02})$$

$$v^2 \omega^2 \rangle = \langle v^2 \rangle \langle \omega^2 \rangle (1 + a_{11})$$



RESULTS: ANGULAR CORRELATIONS



23









9

 $\langle (\widehat{\mathbf{v}}\cdot\widehat{\boldsymbol{\omega}})^2$

$$\langle v^4
angle = rac{5}{3} \langle v^2
angle^2 (1+a_{20})$$

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3b

5

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RESULTS: KURTOSES & SCALAR CORRELATIONS

0.06





$$\begin{split} \langle v^4 \rangle &= \frac{5}{3} \langle v^2 \rangle^2 (1 + a_{20}) \\ \langle \omega^4 \rangle &= \frac{5}{3} \langle \omega^2 \rangle^2 (1 + a_{02}) \\ v^2 \omega^2 \rangle &= \langle v^2 \rangle \langle \omega^2 \rangle (1 + a_{11}) \end{split}$$

CONCLUSIONS AND OUTLOOK

- The Sonine approximation for the temperature ratio successfully corrects the Gaussian prediction.
- While $|a_{20}|$, $|a_{11}|$, and |b| are always small, the cumulant $|a_{02}|$ may take relatively large values, thus invalidating the linear approach (at a quantitative level).
- Comparison with simulation data for the correlation factor b shows that the Sonine approximation with $a_{02}=0$ represents a good compromise between simplicity and accuracy.
- Interesting singular phenomenon in the quasi-elastic limit.
- Simulations planned to test the theoretical predictions.

THANK YOU FOR YOUR ATTENTION!

